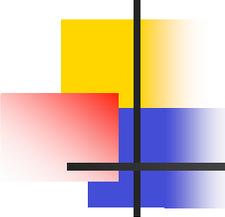


A DARE approach for 3D cloud-resolving simulations of large-scale atmospheric circulation

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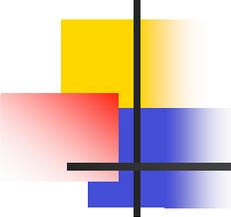


Global CRMs are expensive!

and not everybody has an Earth Simulator...

The Multiscale-Modeling Framework for reducing the computational cost

- Cloud-resolving convective parameterization, or super parameterization
Grabowski, 2001
- A sparse net of 2D CRM (Quasi-3D)
Arakawa, 2004



Diabatic Acceleration and REscaling (DARE)

Kuang, Blossey, Bretherton, GRL, 2005

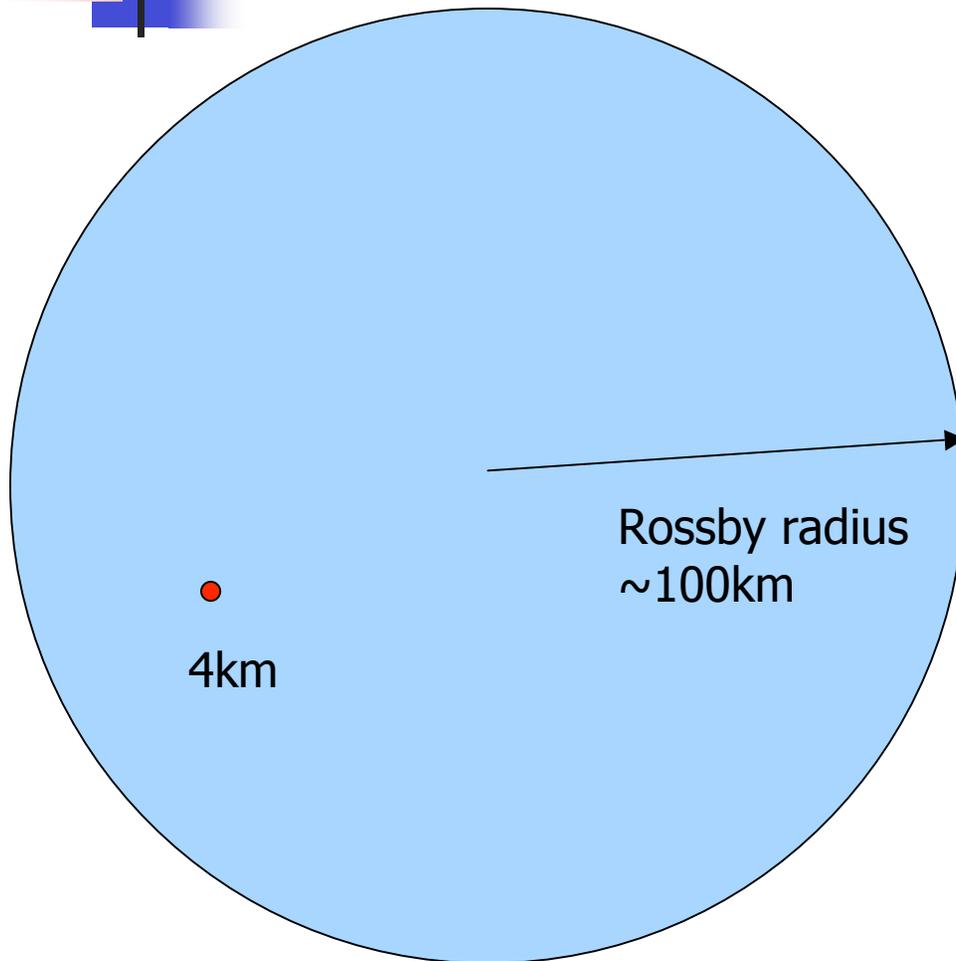
Define the DARE factor $\gamma > 1$

- **1. Accelerate all diabatic processes by a factor of γ**
- **2. Reduce Earth's size and increase its rotation rate by the same factor**

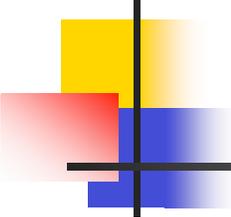
Motivation: reduce the scale separation between the Rossby scale ($\sim 1000\text{km}$) and the cumulus scale (a few km)

A miniature planet $\gamma=10$

(saves a factor of ~ 500 in computation)



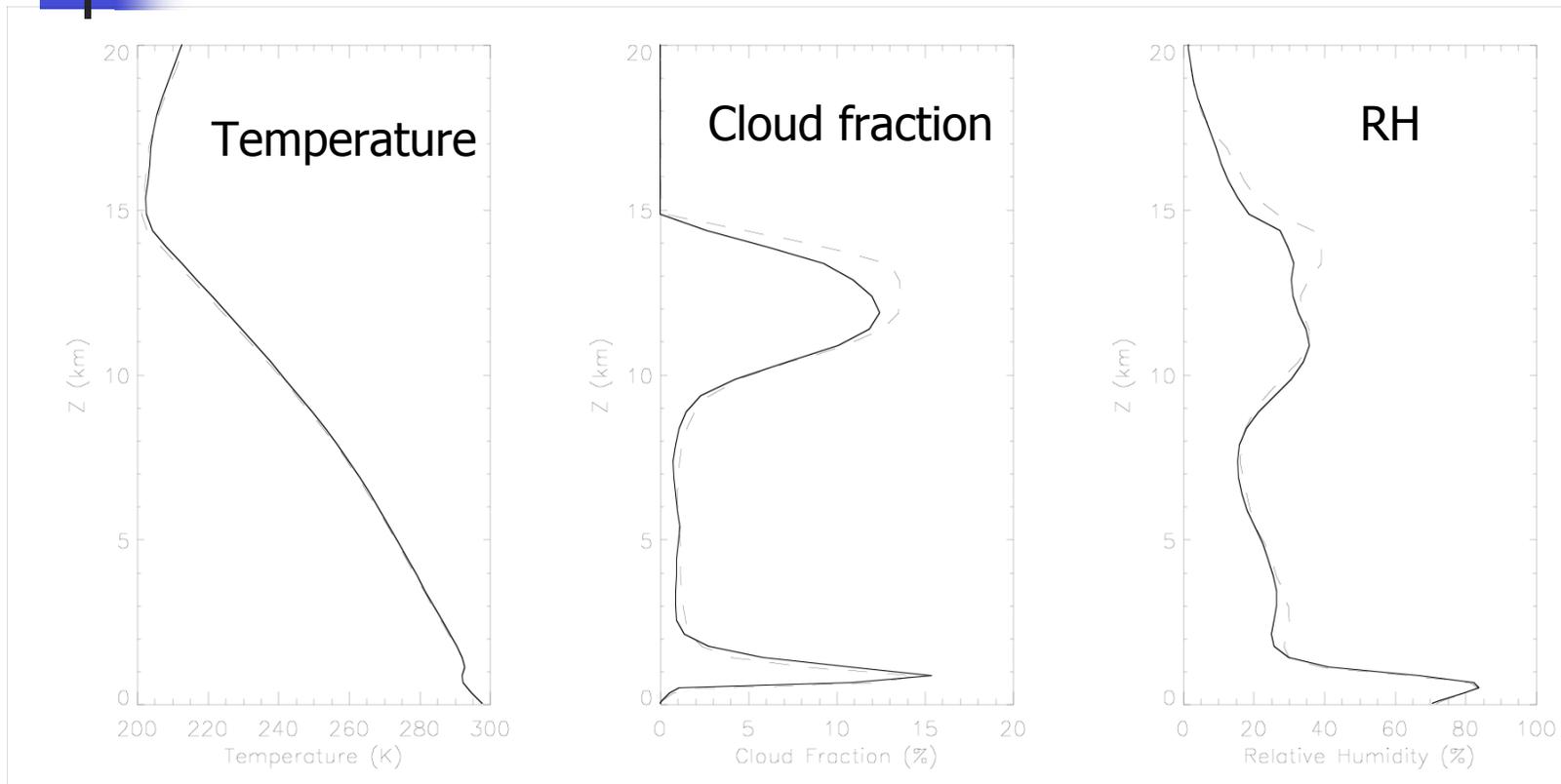
- Assume interactions between the cumulus scale and the Rossby scale remain similar, if the two remain well separated.
- **No artificial scale break**
- Equations for large scale hydrostatically balanced motions unmodified



On the small (convective) scale:

- Turbulence is modeled as usual
- Convection is driven more strongly.
- Can diabatically accelerated convection:
 1. produce reasonable atmospheric profiles?
 2. reproduce the response to large-scale external forcing γ times faster?

Test I: radiative-convective equilibrium profiles



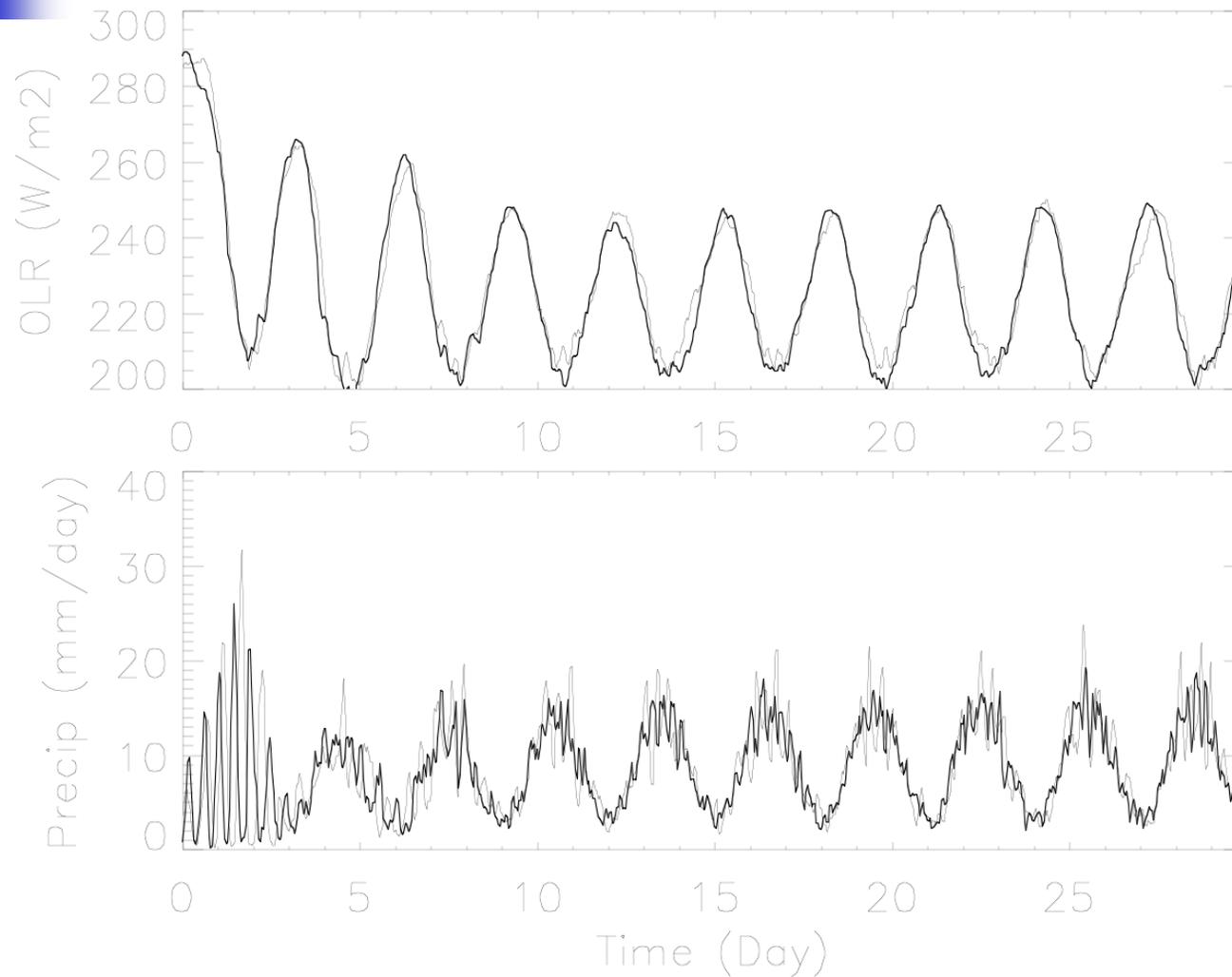
— Control

- - - DARE4

$dx=3\text{km}$, $n_x=n_y=192(\text{control}),48(\text{dare})$
using SAM by Marat Khairoudinov

Test II:

Response to external forcing

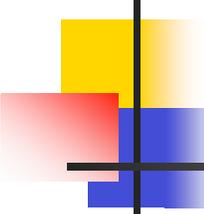


An alternative interpretation:

RAVE (Rescaling of Acceleration in the VERTICAL)

$$\gamma^2 \frac{Dw}{Dt} = -\frac{\partial p'}{\rho \partial z} + B \quad \text{Plus modified SGS model}$$

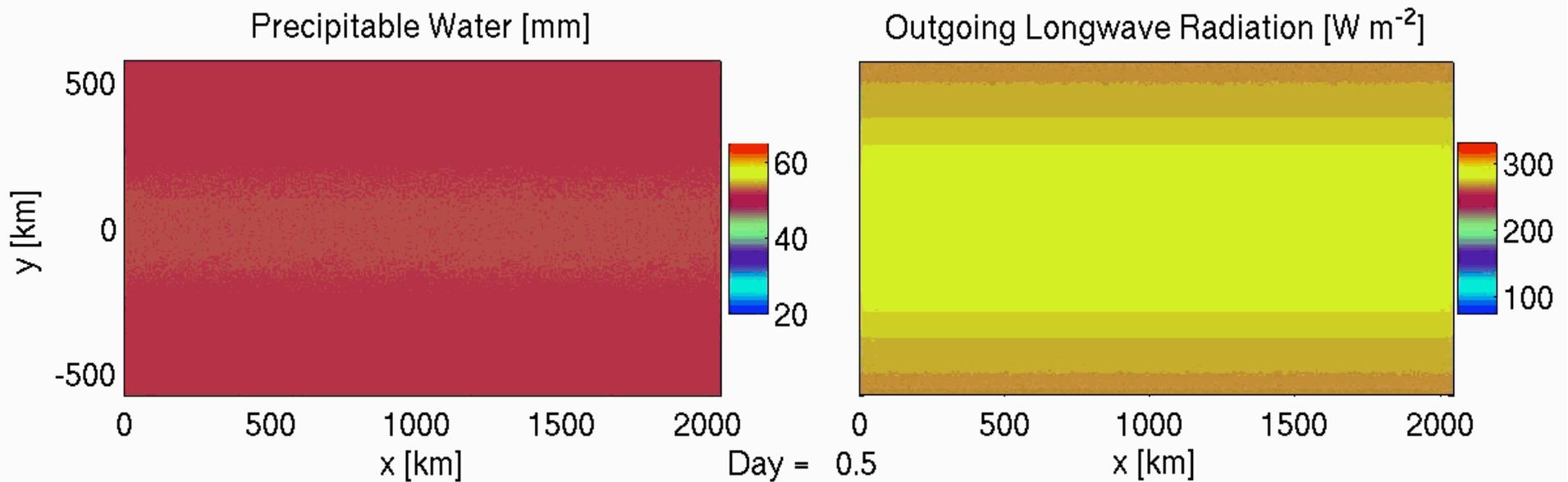
- $\gamma=0$ hydrostatic approximation
 $\gamma=1$ normal, $\gamma>1$: DARE/RAVE
- Hydrostatic motions unmodified
- Can be also viewed as reducing gravity and increasing Z .
$$\frac{D\gamma w}{Dt} = -\frac{\partial p'}{\rho \partial \gamma z} + B / \gamma$$



What is DARE?

- It is one type of coarse horizontal resolution (strongly anisotropic grid) non-hydrostatic model, **so formulated that it behaves like a fine horizontal resolution (more isotropic grid) model that is driven more strongly.**

Equatorial β -plane experiments

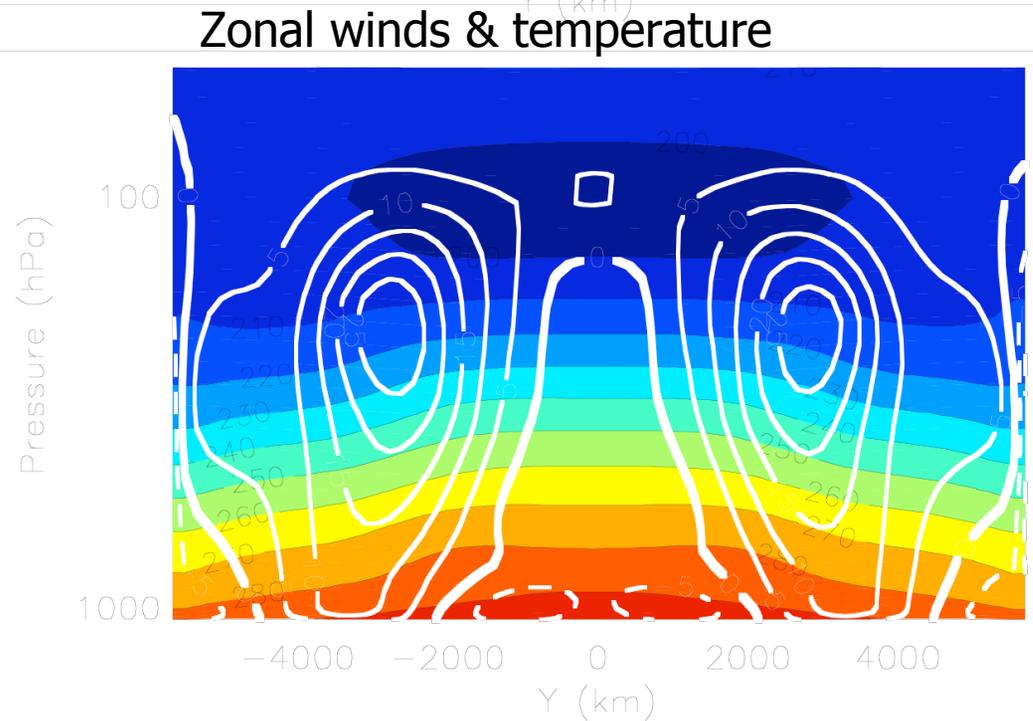
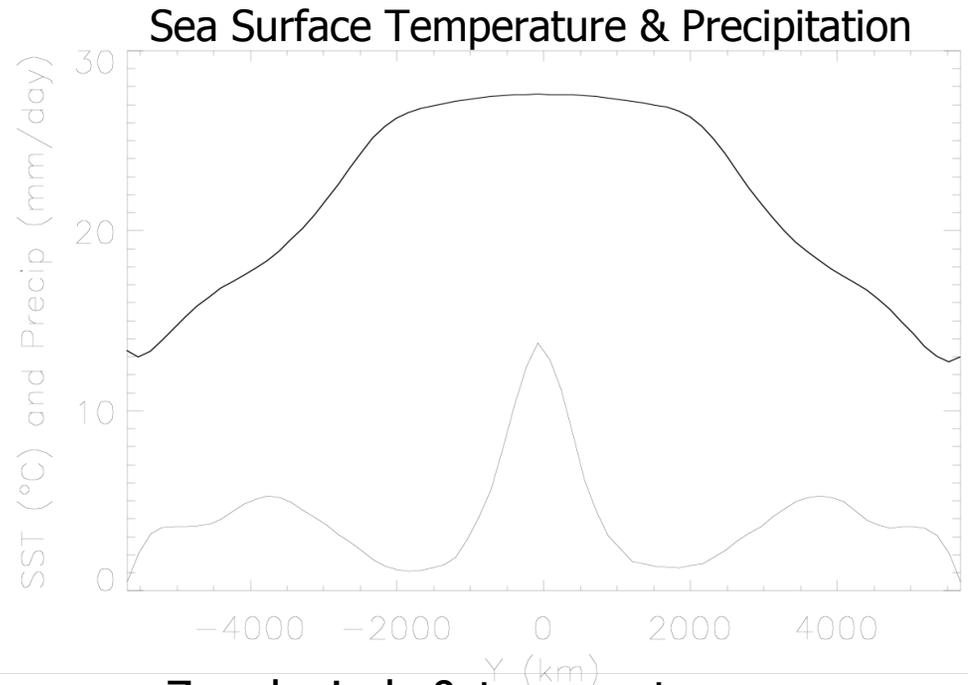
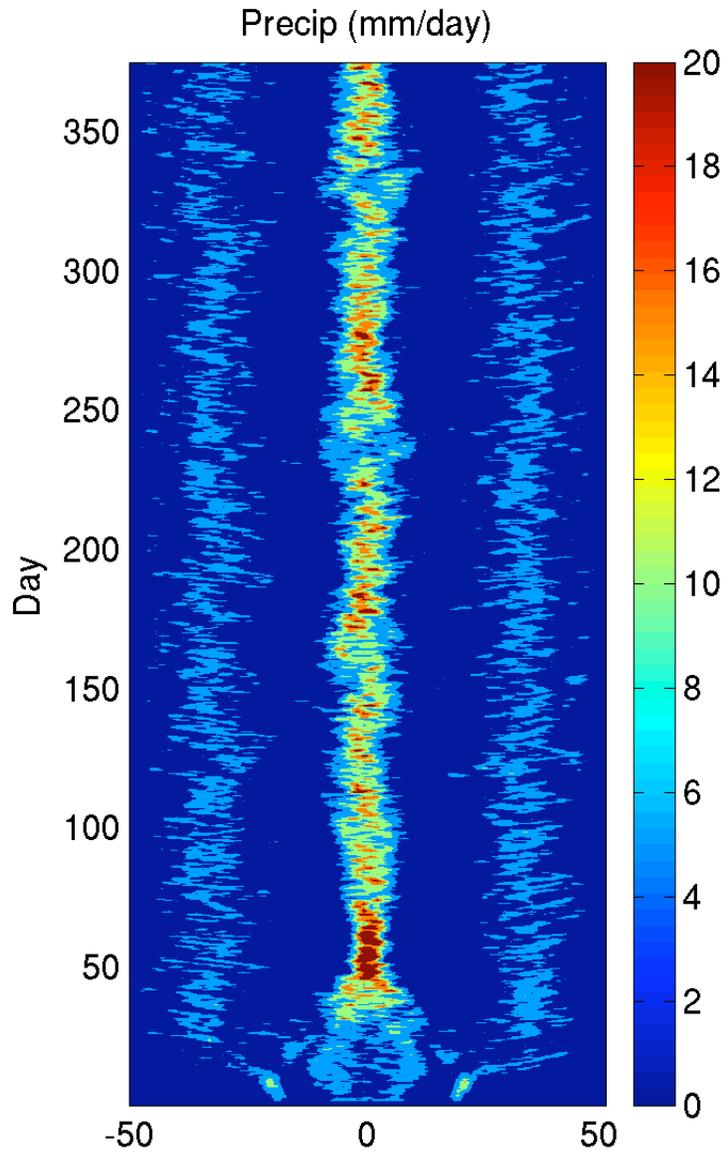


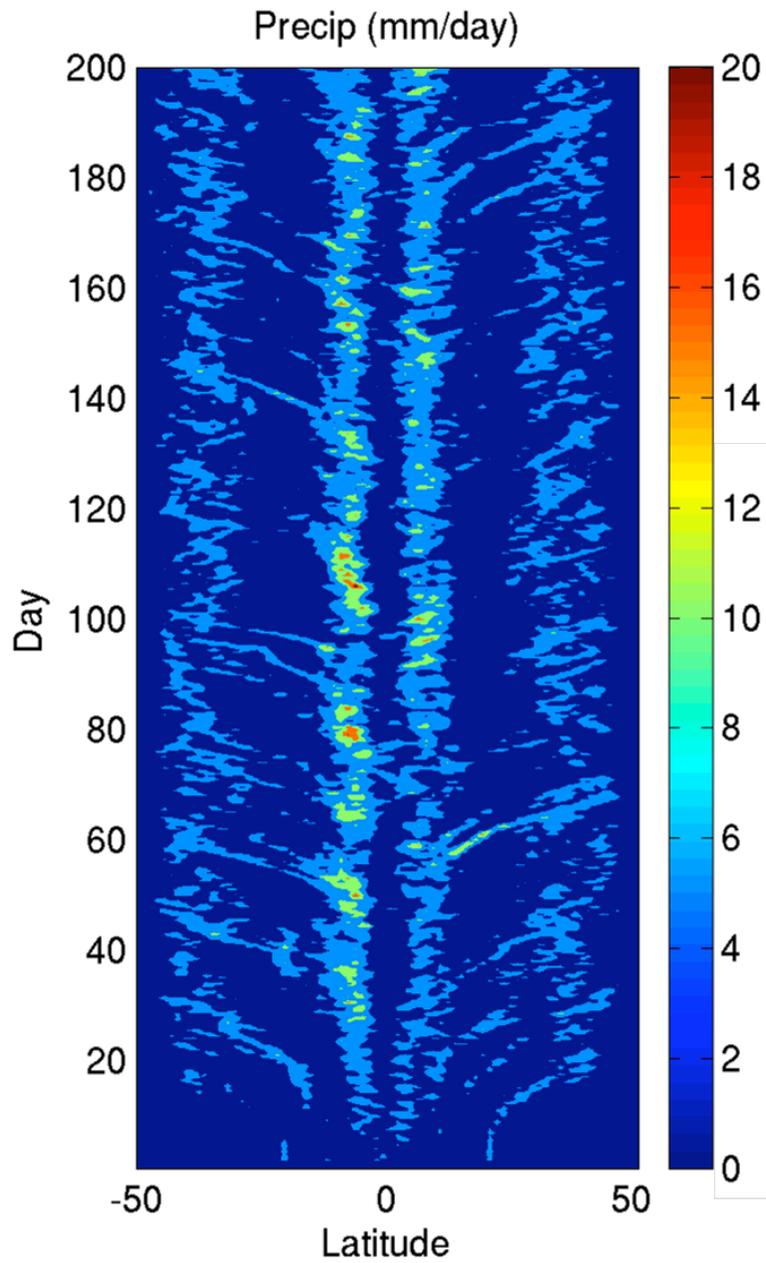
$\gamma=10$, $dx=4\text{km}$, Equinox insolation

15m mixed-layer ocean, zero oceanic heat flux

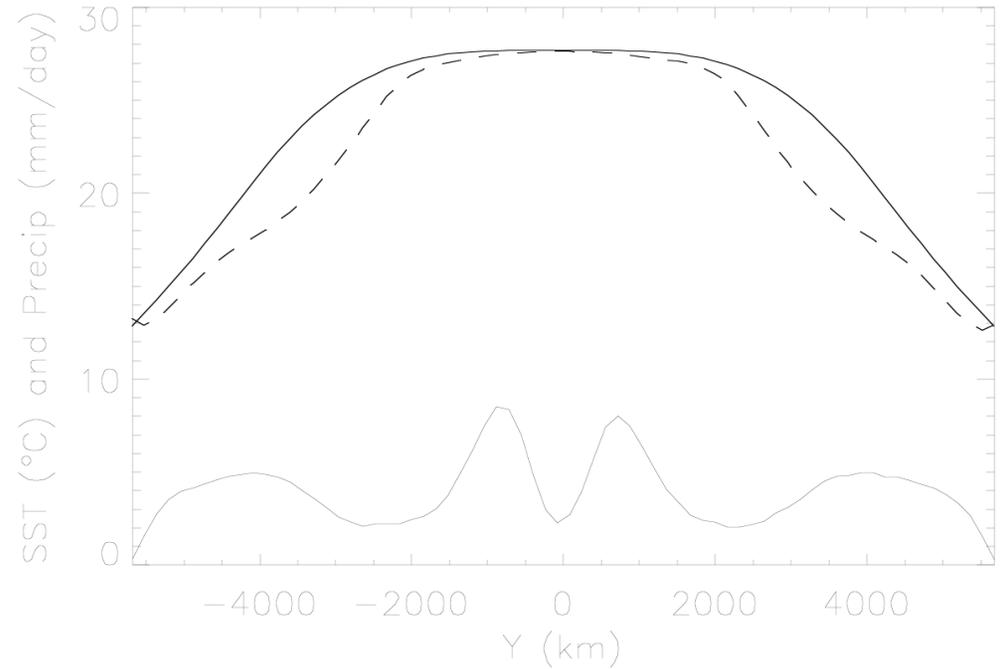
The CRM used is provided by Marat Khairoudinov

A single-ITCZ (the mixed layer ocean case)

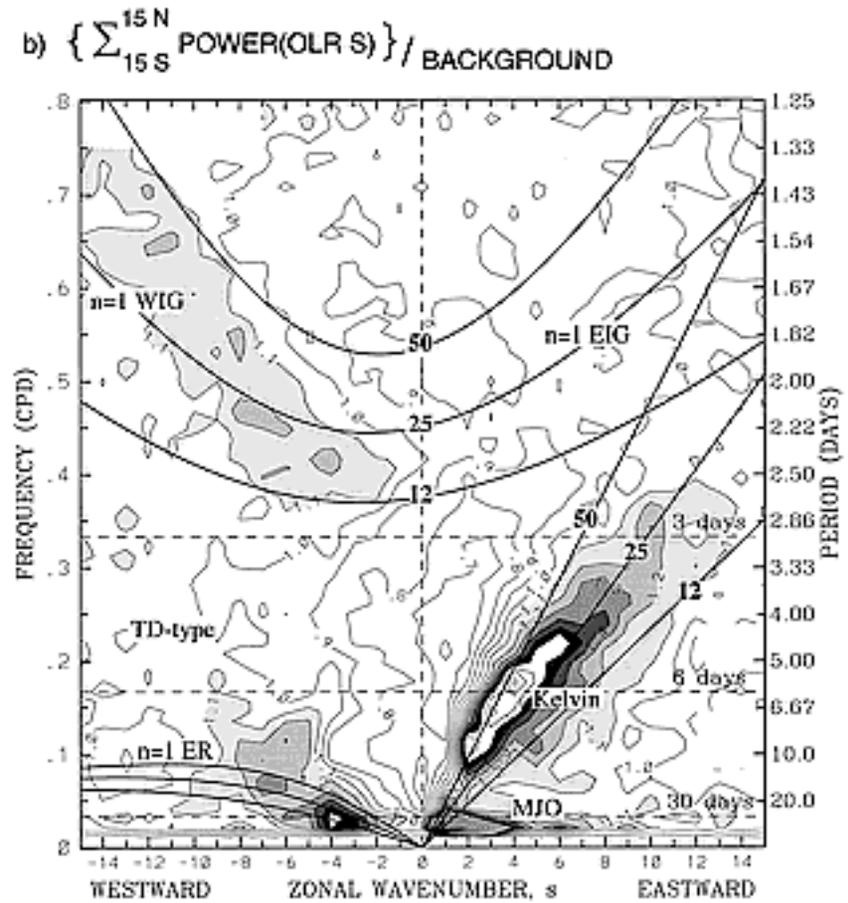
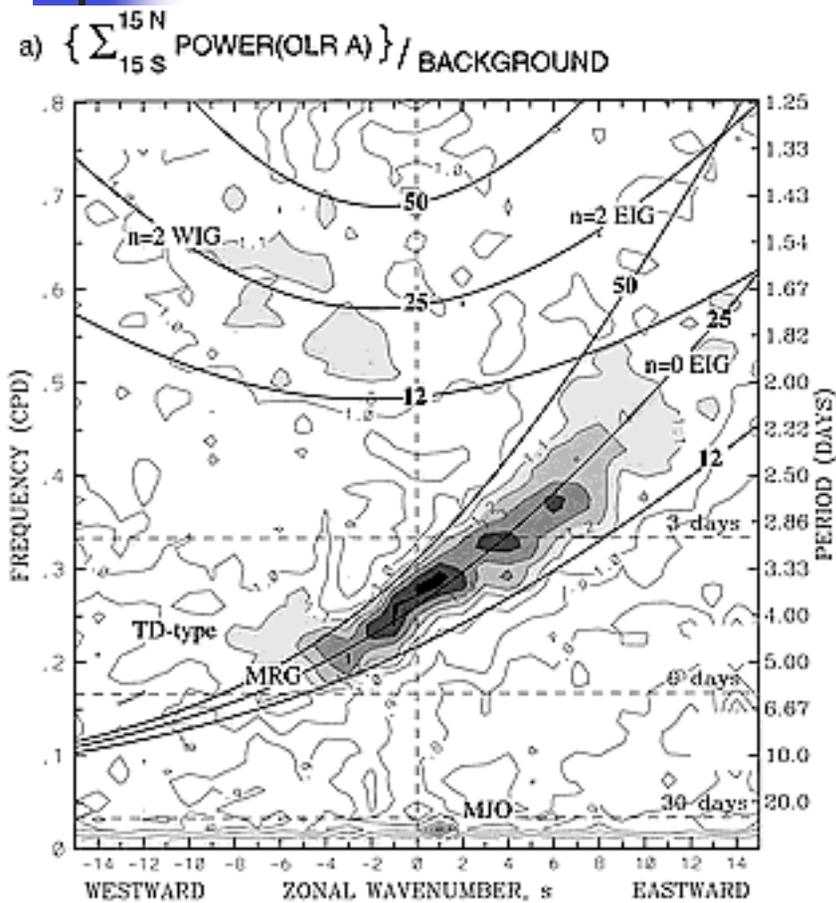




A double-ITCZ
(a fixed SST case)

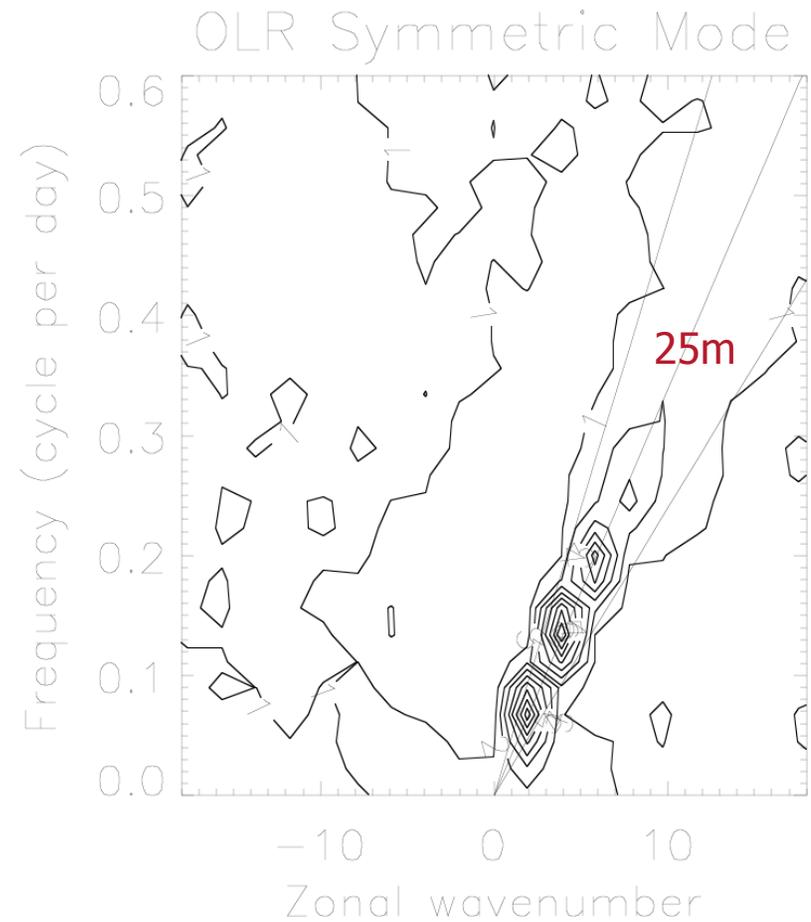
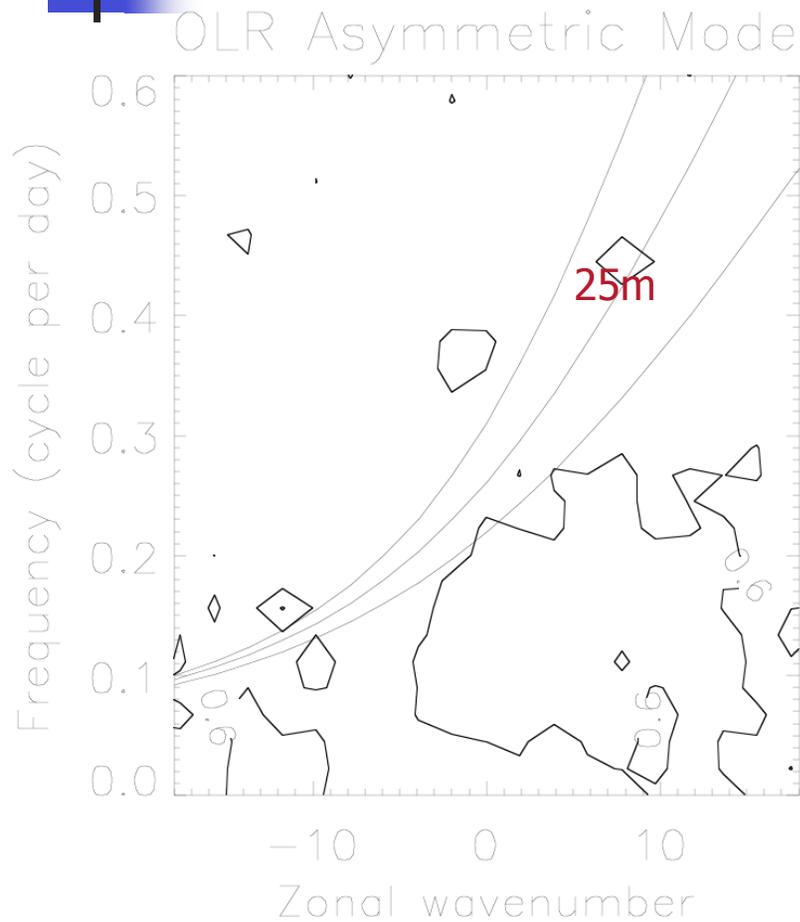


Convectively coupled waves

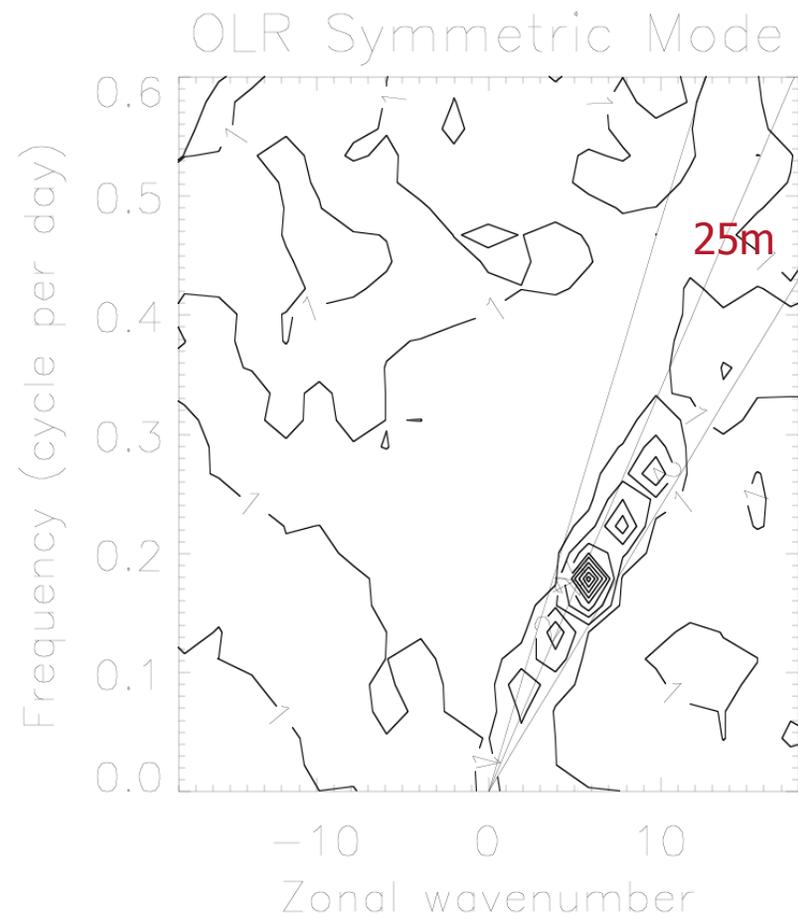
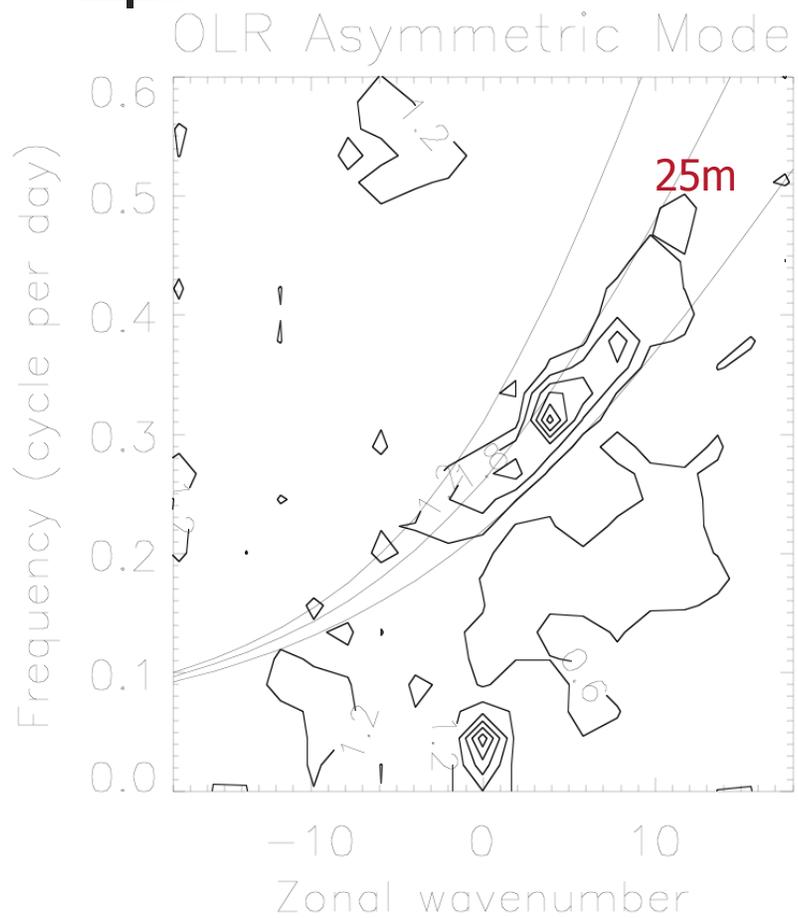


Wheeler and Kiladis, J. Atmos. Sci., 1999

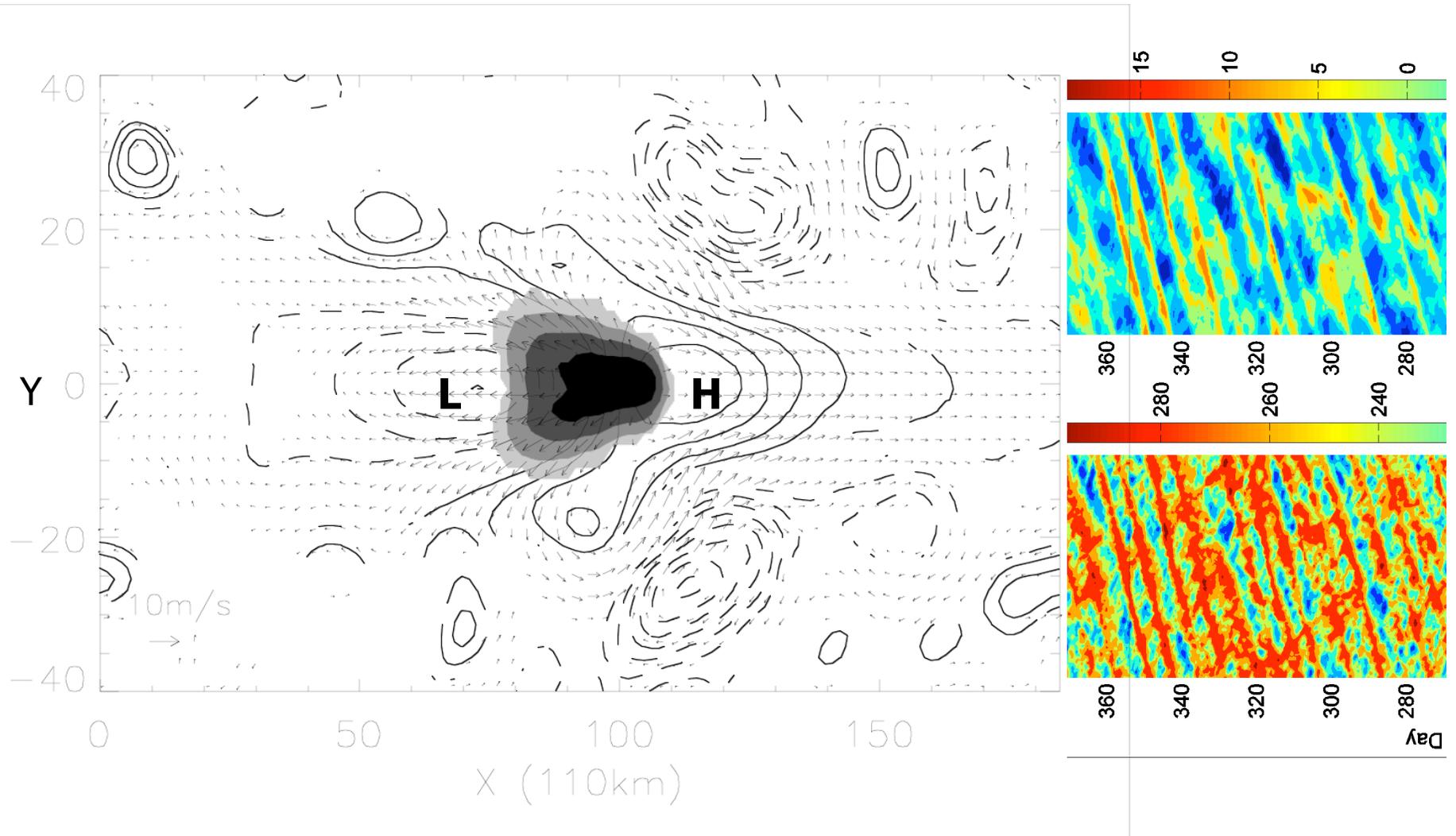
DARE simulation (the single-ITCZ case)



The double-ITCZ case



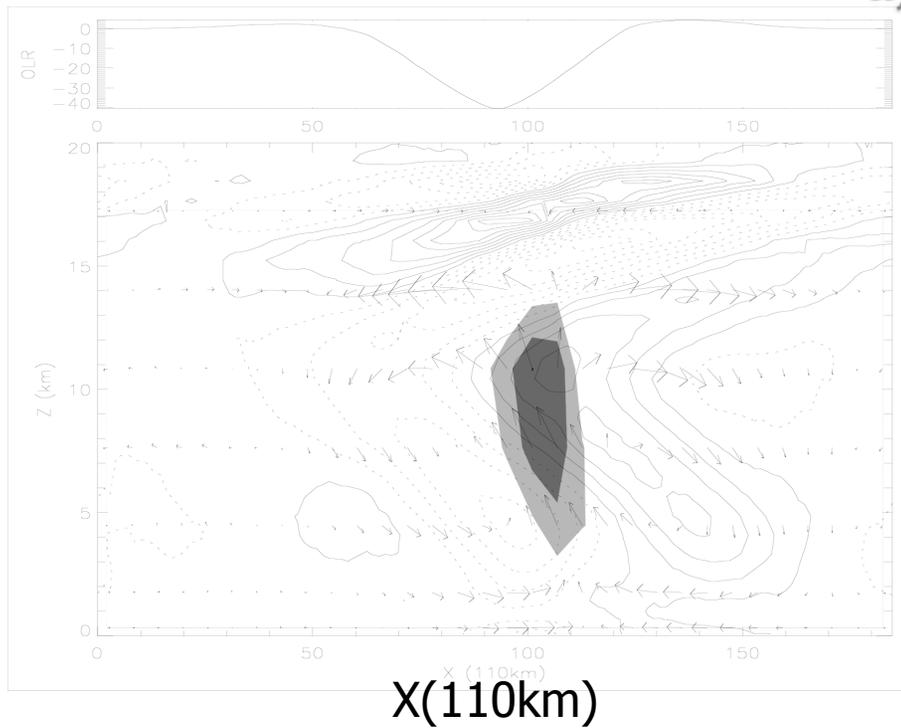
Kelvin waves



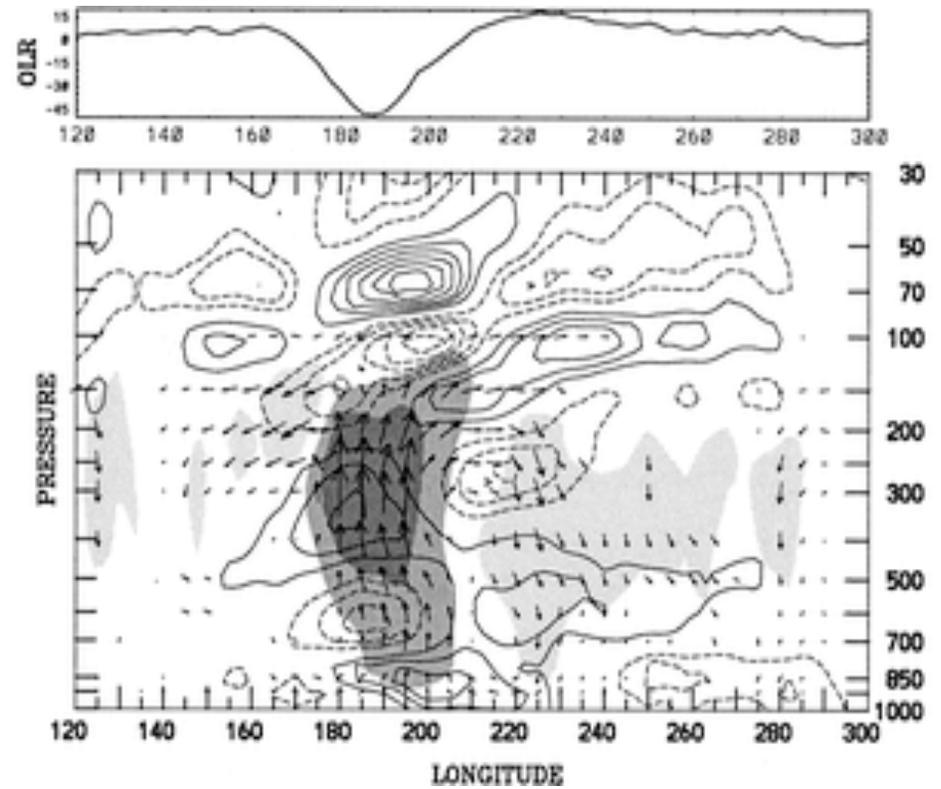
Kelvin waves: vertical structure

DARE temperature & winds

a)



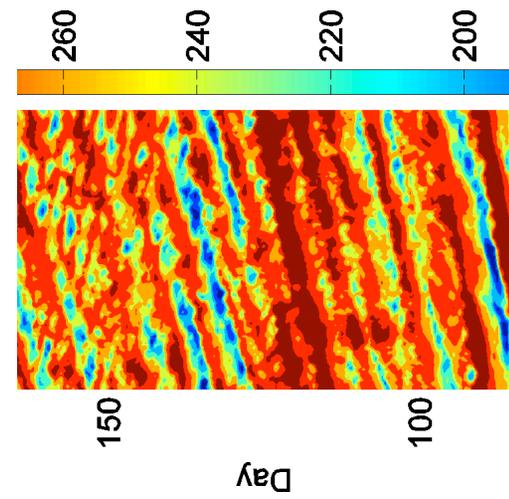
NCEP Reanalysis



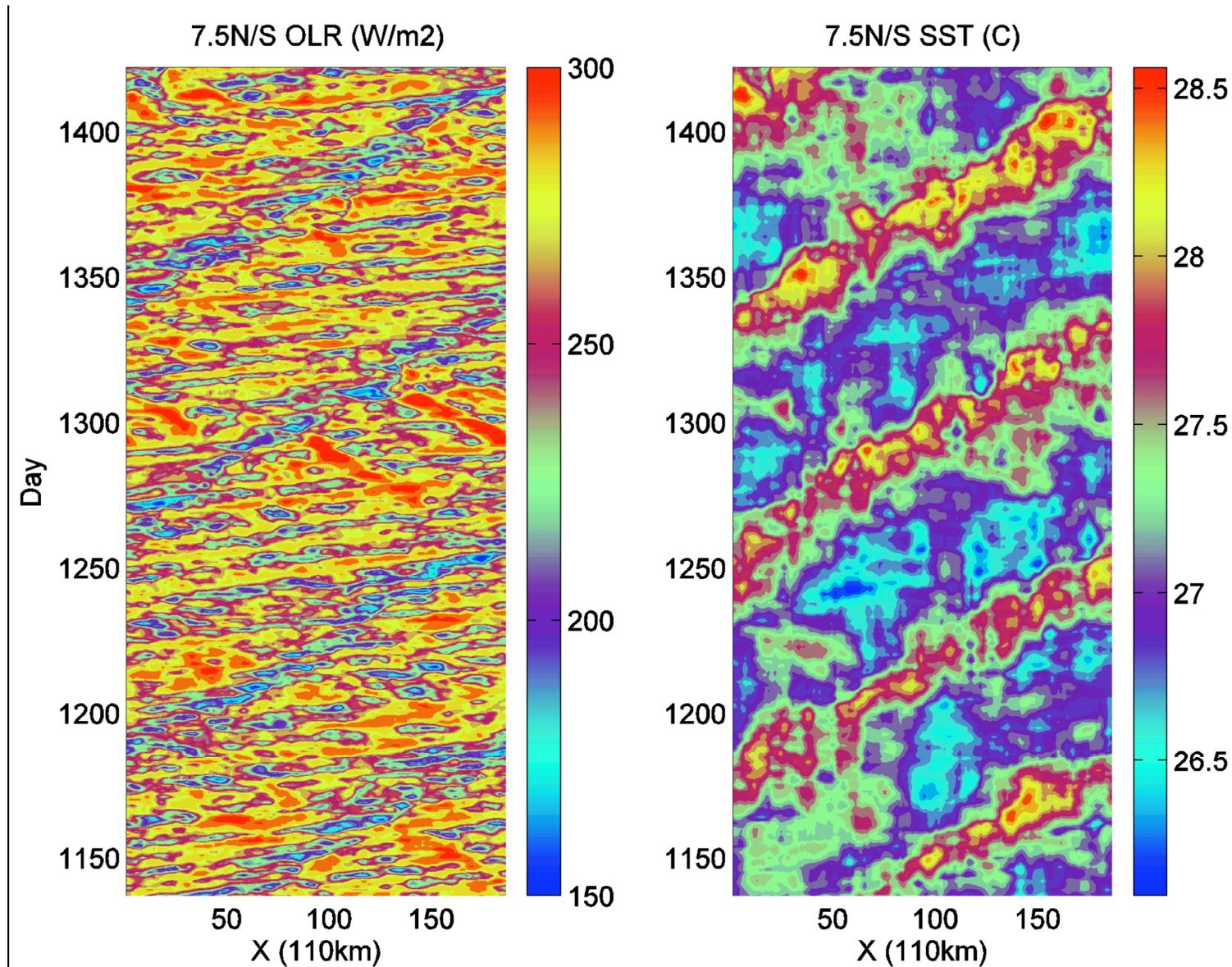
Straub and Kiladis, J. Atmos. Sci., 2003

Testing ideas on the waves

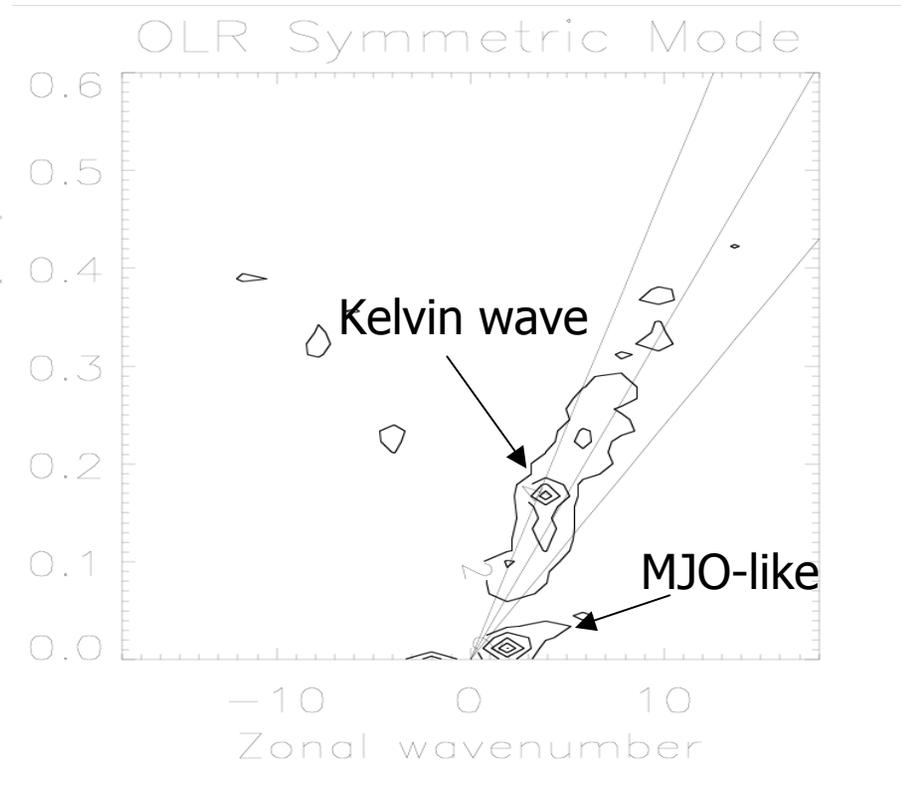
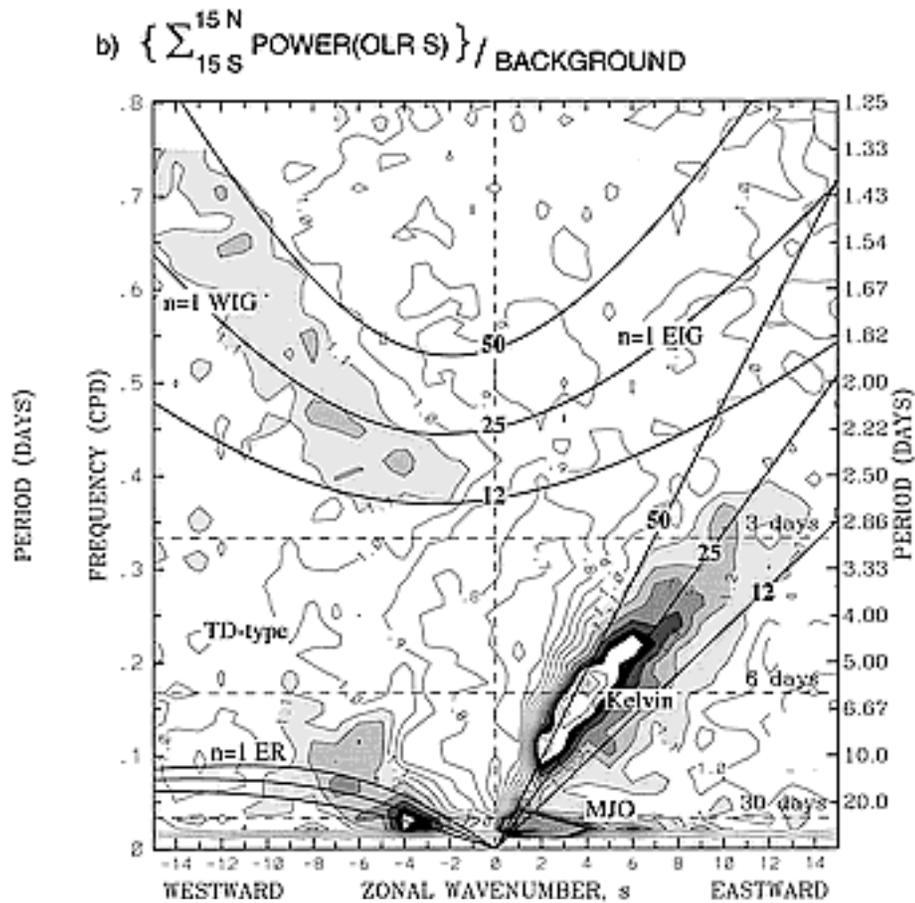
- Kelvin waves remain prominent without WISHE and radiation feedback
- Second baroclinic mode appears essential



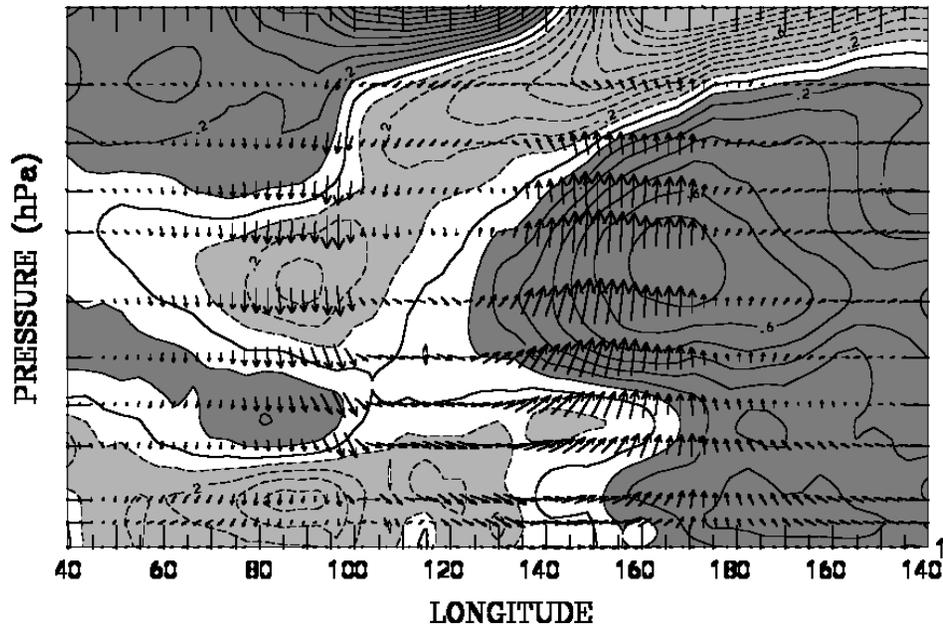
MJO-like disturbances $dx=2\text{km}$ and $\gamma=20$



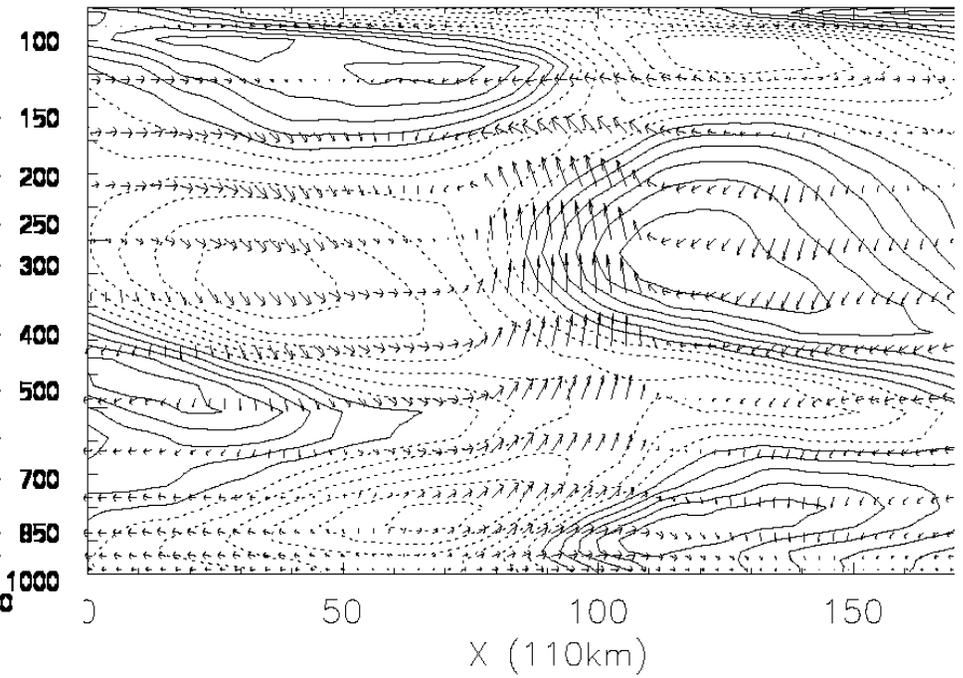
Space-time spectrum



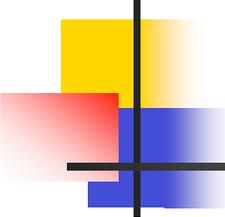
“MJO” Structure



Observed, Kiladis et al, 2005

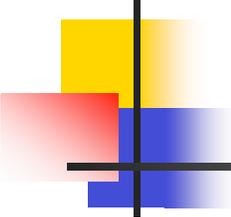


DARE simulation



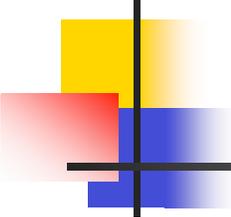
Conclusions

- We propose DARE as a way to simulate the interaction between large-scale circulation and convection
- Easy to implement; fully 3D; fine with land/topography; saving in computation $\sim O(\gamma^3)$;
- As γ reduces to 1, the DARE model smoothly converges to a global CRM
- Can be used within MMF as well



Conclusions (cont.)

- The simulated convectively coupled waves and “MJO” resemble those observed in terms of their spectra and structures.
- Convective coupled waves exist without WISHE or radiative feedback
- The simulated “MJO” only exists with ocean coupling and surface/radiative flux feedbacks



Ongoing work

- Better understand the effect of DARE and explore potential improvements
- Include shallow convective parameterization
- Examine mechanisms of convectively coupled waves, and realism of the simulated MJO, its sensitivity to resolution, microphysics/ radiation schemes etc...
- Realistic climate/weather simulations with DARE