

# Bringing climate physics (better radiation) to cloud models and the MMF

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# Radiation is a weak link in the MMF and in SAM

SAM's current radiation package (optics + solver) comes from CAM 3. It's no longer state-of-the-art, especially for absorption by gases, and NCAR is replacing it with RRTMG

Even the old radiation package is expensive, and we take shortcuts: radiation is computed every  $N$  times compared to other processes

I implemented RRTMG (the next-generation scheme for CAM) in SAM

Oops - it's about six times slower. That won't work

About this time Bjorn called me with the same problem in LES simulations of stratocumulus using a different (and very streamlined) code

# Broadband spectral integration is the issue

Radiative transfer is fundamentally non-linear, and properties vary with wavelength

RRTMG uses a  $k$ -distribution

$$F(i, j, t) = \sum_b^B w_b \sum_g^{G(b)} w_g F_{b,g}(i, j, t)$$

Effects of clouds, etc. →

Effects of gases →

Flux per g-point ←

In RRTMG,  $B$  and  $G$  are both about 16 - flux estimates take 240 calculations

# Can we possibly afford realistic radiation?

Reducing the frequency of radiation calculations is a bad idea for practical and conceptual reasons

Accuracy in radiation don't degrade gracefully in number of bands/g-points

We had an insight that amounts to computational judo



## What if...

Doing the entire calculation every  $N$  time steps costs the same as doing  $1/N$  calculations every time step.

We took this idea to the extreme

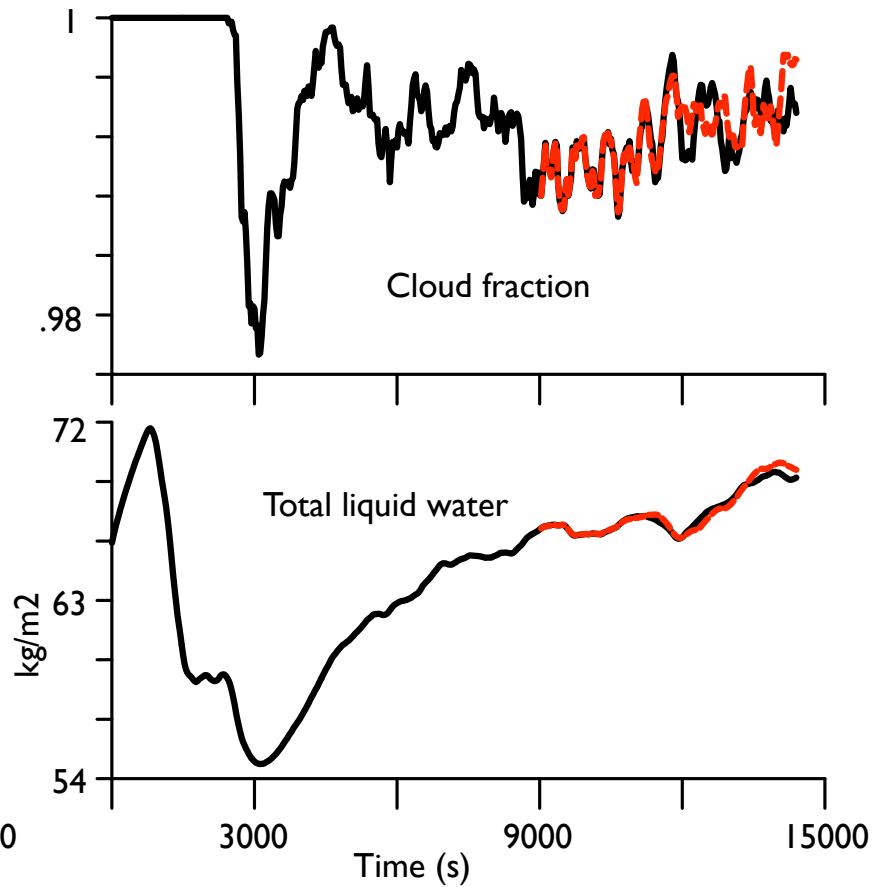
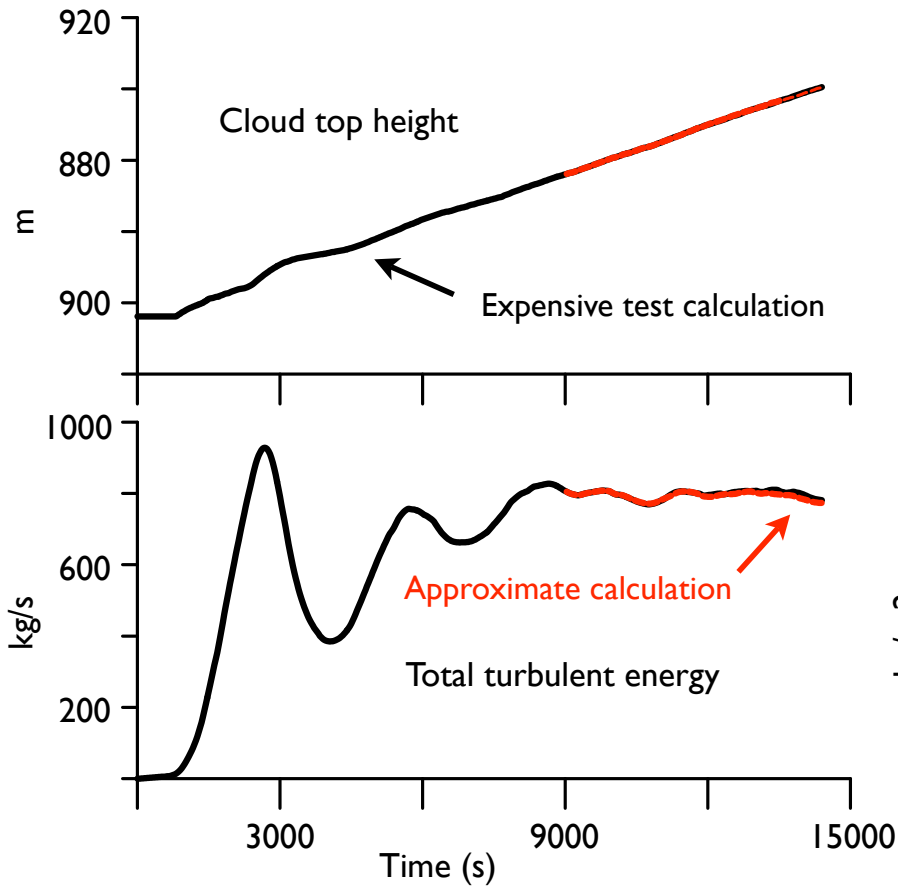
$$F(i, j, t) = \sum_b^B w_b \sum_g^{G(b)} w_g F_{b,g}(i, j, t)$$

becomes

$$F(i, j, t) \approx w_{b'} F_{b',g'}(i, j, t)$$

where  $b', g'$  pair is chosen randomly at each grid point and done every time step

# We built this. It works like a champ.



# How this works

This trick spreads out the correct calculation in time and space

The sampling introduces lots of noise but *no systematic errors* - so the circulation isn't changed  
("Heating rate errors are orthogonal to the flow.")

Formally, this is a Monte Carlo sample of the calculation we'd like to do but can't afford, so...

"Monte Carlo spectral integration"



# Why does it work?

Uncorrelated noise can't excite organized motions  
(c.f. Shutts stochastic parameterizations)

We did a scaling analysis for well-mixed stratocumulus  
comparing  
the energy excited by noise at a given time & space scale  
with  
the energy expected from organized flow  
to demonstrate this

The idea should work just as well for cloud resolving  
models (SAM, VVM)



# What's next?

We now have a way to make state-of-the-art radiative transfer calculations in cloud-scale models

We needed this for the MMF but it's very good for cloud models in general

We also have a theoretical framework to assess other approximations (especially 3D vs 1D radiative transfer)

We're moving the idea into the MMF in stages (implementation in RRTMG, tests with stand-alone SAM, then testing the full MMF)