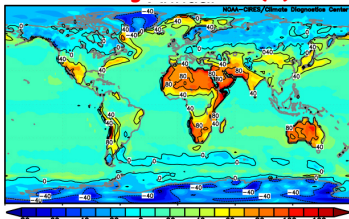
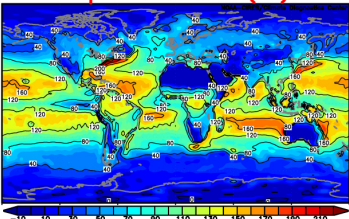


Energy from the Surface to the Air

Rising Warm Air (H)

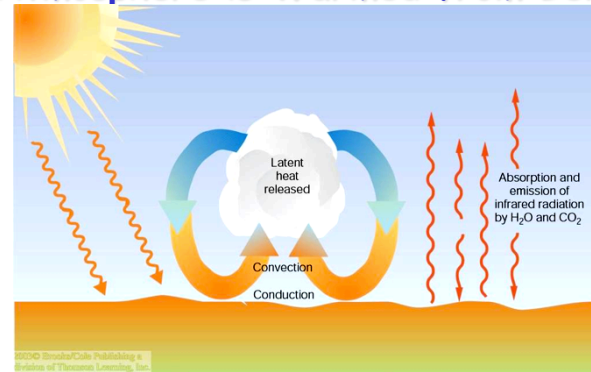


Evaporated Water (LE)



- Energy absorbed at the surface warms the air
- Some of this energy is transferred in rising warm “thermals”
- But more of it is “hidden” in water vapor

Atmosphere is Warmed from Below

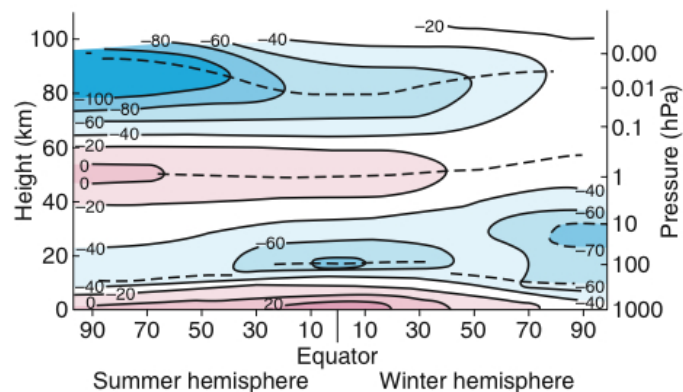


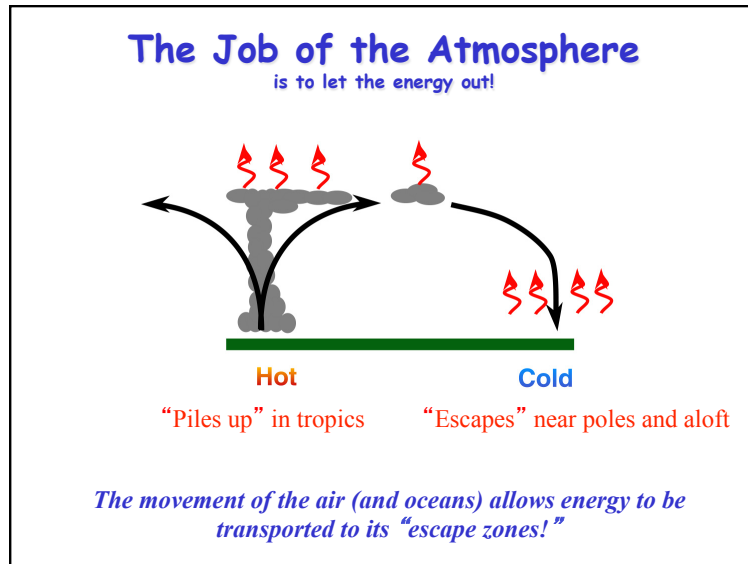
Solar radiation passes first through the upper atmosphere, but only after absorption by earth's surface does it generate sensible heat to warm the ground and generate longwave energy. This heat and energy at the surface then warms the atmosphere from below.

Vertical Structure is Crucial

- The world is a big place, but the **atmosphere is very thin**, and most of it is close to the ground
 - About **15% of the atmosphere is below our feet**
 - At the top of Long's Peak, the figure is 40%
 - You are closer to outer space than you are to Denver!
- Changes in atmospheric temperature with height are responsible for the **“Greenhouse Effect,”** which keeps us from freezing to death

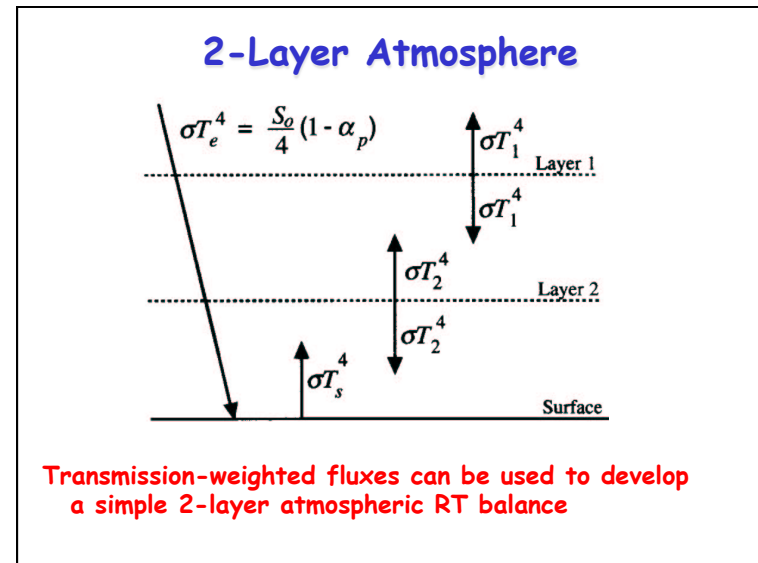
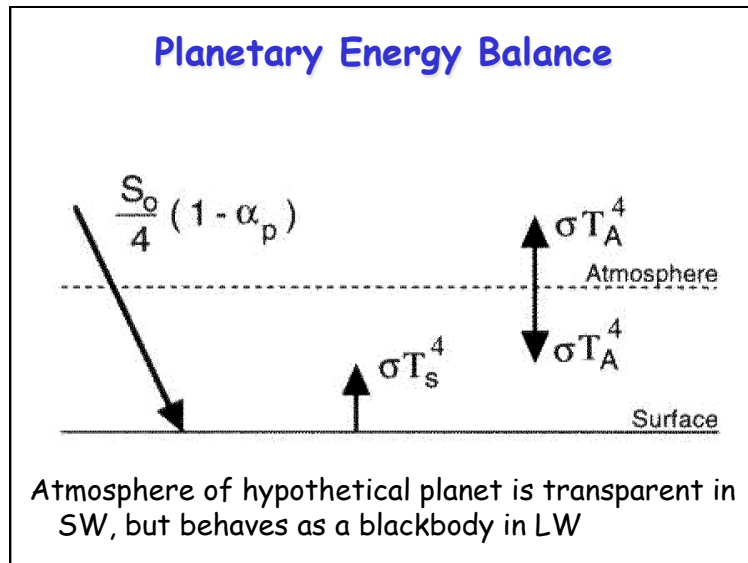
Vertical Structure of the Air





Atmospheric Heating by Convection

- Sunlight warms the ground
- Ground warms adjacent air by conduction
 - Poor thermal conductivity of air restricts heating to a few cm
- Hot air forms rising air "bubbles" (thermals) leading to convection ... heats the air, but cools the surface!
 - Mechanical mixing due to wind enhances this mode of heat transport



Radiative Balances by Layer

For every layer:
Energy In = Energy Out

TOA $\frac{S_0}{4}(1-\alpha_p) = \sigma T_1^4$

L1 $\sigma T_2^4 = 2\sigma T_1^4$

L2 $\sigma T_s^4 + \sigma T_1^4 = 2\sigma T_2^4$

Surface $\frac{S_0}{4}(1-\alpha_p) + \sigma T_2^4 = 2\sigma T_s^4$

2-Layer BB Atmosphere (cont' d)

Vertical temperature profile for 2-layer atmosphere, with thin graybody layers at top and bottom.

**Very unrealistic lapse rate!!
Why?**

- Solving energy budgets for all layers simultaneously gives

$$T_s^4 = 3 \frac{(S_0/4)(1-\alpha_p)}{\sigma} = 3T_e^4$$

- Recall from Ch 2 that a 1 layer B-B atmosphere produces $T_s^4 = 2T_e^4$
- In general, an n-layer B-B atmosphere will have $T_s^4 = (n+1)T_e^4$

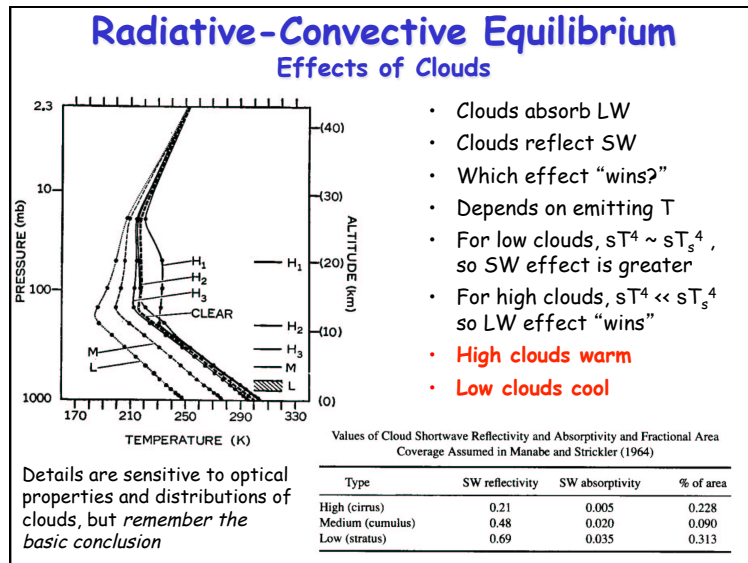
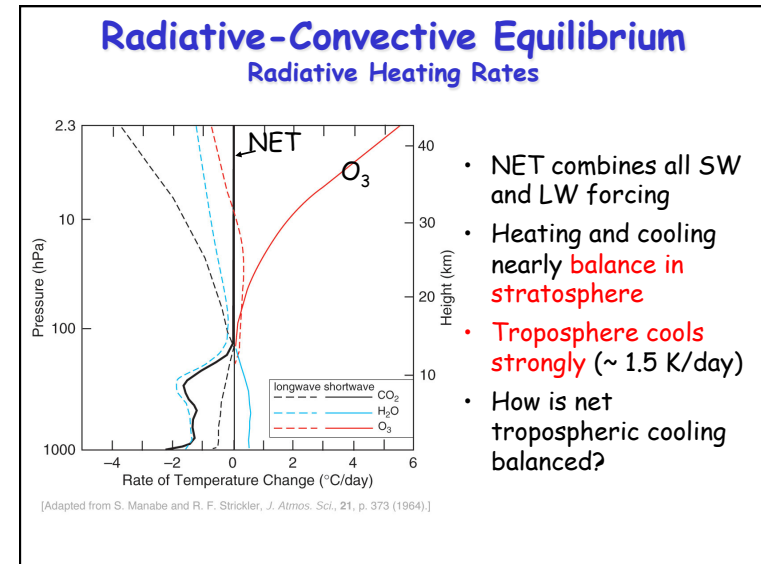
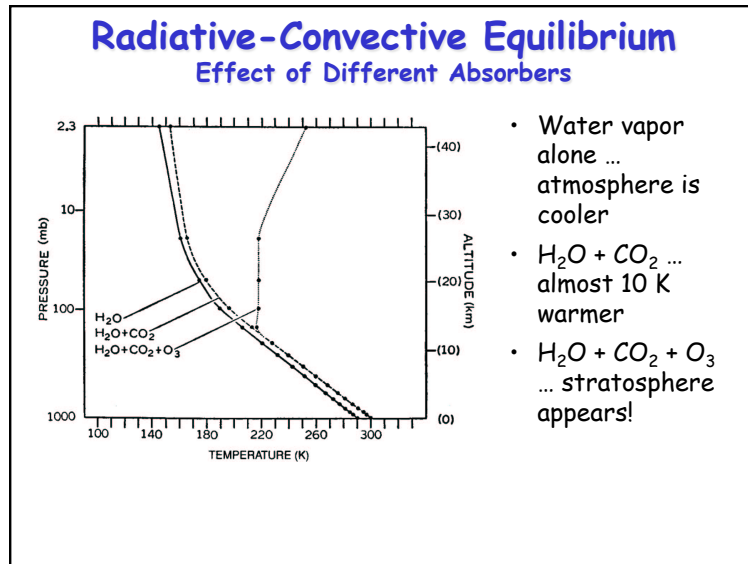
Radiative-Convective Models (a recipe)

- Consider a 1-D atmosphere
- Specify solar radiation at the top, emissivity of each layer
- Calculate **radiative equilibrium temperature** for each layer
- Check for **static stability**
- If layers are **unstable, mix them!**
 - (e.g. if $G > G_d$, set both T 's to mass-weighted mean of the layer pair)
- Add **clouds and absorbing gases** to taste

Manabe and Strickler (1964)

Radiative-Convective Equilibrium

- Pure radiative equilibrium is **way too hot** at surface
- Adjusting to G_d still too steep
- Adjusting to observed 6.5 K km^{-1} produces fairly reasonable profile:
 - Sfc temp (still hot)
 - Tropopause (OK)
 - Stratosphere (OK)



Global Mean Cloud Radiative Forcing

	Average	Cloud-free	Cloud forcing
OLR	234	266	+31
Absorbed solar radiation	239	288	-48
Net radiation	+5	+22	-17
Albedo	30%	15%	+15%

Radiative flux densities are given in $W m^{-2}$ and albedo in percent. [From Harrison *et al.* (1990), © American Geophysical Union.]

- Clouds increase planetary albedo from 15% to 30%
- This reduces absorbed solar by $48 W m^{-2}$
- Reduced solar is offset by $31 W m^{-2}$ of LW warming (greenhouse)
- So total cloud forcing is $-17 W m^{-2}$
- Clouds cool the climate. By how much? How might this number change if cloudiness increased?

